Control of attosecond light polarization in two-color bicircular fields

Á. Jiménez-Galán, ^{1,*} N. Zhavoronkov, ¹ D. Ayuso, ¹ F. Morales, ¹ S. Patchkovskii, ¹ M. Schloz, ¹ E. Pisanty, ^{1,2} O. Smirnova, 1,3 and M. Ivanov 1,4,5

¹Max-Born Institute for Nonlinear Optics and Short Pulse Spectroscopy, Max-Born-Straße 2A, D-12489 Berlin, Germany ²ICFO-Institut de Ciències Fotòniques, The Barcelona Institute of Science and Technology, 08860 Castelldefels (Barcelona), Spain ³Technische Universität Berlin, Ernst-Ruska-Gebäude, Hardenbergstraße 36A, 10623 Berlin, Germany ⁴Department of Physics, Imperial College London, South Kensington Campus, SW72AZ London, United Kingdom ⁵Institute für Physik, Humboldt-Universität zu Berlin, Newtonstraße 15, D-12489 Berlin, Germany



(Received 17 October 2017; published 8 February 2018)

We develop theoretically and confirm both numerically and experimentally a comprehensive analytical model which describes the propensity rules in the emission of circularly polarized high harmonics by systems driven by two-color counter-rotating fields, a fundamental and its second harmonic. We identify and confirm the three propensity rules responsible for the contrast between the 3N + 1 and 3N + 2 harmonic lines in the HHG spectra of noble gas atoms. We demonstrate how these rules depend on the laser parameters and how they can be used in the experiment to shape the polarization properties of the emitted attosecond pulses.

DOI: 10.1103/PhysRevA.97.023409

I. INTRODUCTION

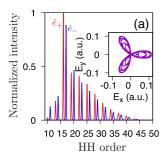
Circular or highly elliptical light pulses in the extreme ultraviolet spectral range offer numerous applications in chiralsensitive light-matter interactions in both gas and condensed phase, ranging from chiral recognition [1,2] to time-resolved magnetization dynamics and spin currents [3–9], not only in condensed matter but also in isolated atoms [10]. Until recently, such radiation has only been available at large-scale facilities (e.g., synchrotrons, XFELs), with ultrafast time resolution requiring free electron lasers. On the other hand, laboratorybased sources of pulses of attosecond duration (1 as = 10^{-18} s) are becoming broadly available, allowing one to monitor and control electronic dynamics at their intrinsic (attosecond) time scales [11-17]. However, the available attosecond pulses are typically linearly polarized, which limits their applicability making them unable to probe and control processes such as ultrafast spin and chiral dynamics. These latter require the generation of attosecond pulses with controllable degrees of polarization.

Ultrashort extreme-ultraviolet (XUV) or x-ray pulses are obtained via the high harmonic generation (HHG) process [15] and their production is linked to the recombination of the electron promoted to the continuum with the hole it has left in the parent atom or molecule [18]. Such recombination is more likely when the electron is driven by a linearly polarized laser field, leading to linearly polarized harmonics. The suppression of electron return to the parent ion in strongly elliptic fields means that brute-force approach to the generation of highly elliptically polarized attosecond pulses by using a highly elliptical driver fails. One practical solution to this problem is to use two-color driving fields in the so-called bicircular configuration, originally proposed in [19-21]. Following elegant recent experiments [22], this proposal has finally attracted the attention it has deserved [22-31], including schemes for the generation of chiral attosecond pulses [26,32–36].

The bicircular configuration consists of combining a circularly polarized fundamental field with its counter-rotating second harmonic. The Lissajous figure of the resulting field is shown in the inset of Fig. 1(a). It is symmetric with respect to a rotation of 120°. Each of the leaves in the trefoil generates an attosecond burst, totalling three bursts per laser cycle. In the frequency domain, this field produces circularly polarized harmonic peaks at $(3N + 1)\omega_{IR}$ and $(3N + 2)\omega_{IR}$ values, with the helicity of the fundamental field and the second harmonic, respectively, while $3N\omega_{\rm IR}$ harmonic lines are symmetry forbidden (see, e.g. [21–27,30,37]).

The alternating helicity of the allowed harmonics is, however, an important obstacle en route to producing circularly polarized attosecond pulses or pulse trains. Indeed, to achieve this goal one set of the harmonic lines, e.g., 3N + 1, with well-defined helicity, must dominate over the other set (e.g., 3N + 2), across a wide range of spectral energies. This is not always the case; see, e.g., Refs. [19,36] for measurements in argon or helium or Fig. 1(a) for calculations in helium. Kfir and collaborators reported substantial suppression of the (3N + 2)harmonic lines of neon by optimizing phase-matching conditions in a gas-filled hollow fiber [27,32]. Further analysis demonstrated that such suppression appears already at the microscopic level when the system is ionized from the 2p orbital of neon [see Fig. 1(b)] [26,36,38], opening the way to practical generation of polarization-controlled attosecond bursts. Recently, Ref. [39] reported that changing the intensity ratio between the two circular drivers offers a possibility to suppress even further the 3N + 2 lines. The exact physical origins of the suppression at the atomic level and the physical mechanisms responsible for it are, however, not completely understood. While Ref. [26] argued that the origin of the suppression is linked to the initial angular momentum of the

^{*}jimenez@mbi-berlin.de



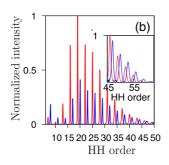


FIG. 1. Numerical TDSE HHG spectrum of (a) helium and (b) neon. The inset in (a) shows the y vs x components of the driving electric field. For helium, the fundamental and second harmonic drivers have field strengths of $F_{\omega}=F_{2\omega}=0.056$ a.u. and duration of 12 fs. For neon, the field strengths of the drivers are $F_{\omega}=F_{2\omega}=0.073$ a.u. and their duration is 12 fs. The red lines indicate the $\hat{\mathbf{e}}_+$ spherical component, which corresponds to light rotating with the fundamental driver (counterclockwise), while the blue line indicates the $\hat{\mathbf{e}}_-$ spherical component, which corresponds to light rotating with the second harmonic (clockwise). Angular momentum conservation imposes that the 3N+1 harmonic lines rotate with the fundamental field, the 3N+2 lines rotate with the second harmonic, and the 3N lines are forbidden. The inset in (b) shows the cutoff harmonics for which the blue lines start to dominate in the spectrum of neon.

ionizing orbital, Ref. [36] suggested the Cooper-like suppression of the recombination step to be chiefly responsible.

Here, we first present experimental results for neon, showing control over the contrast between the 3N+1 and 3N+2 harmonic lines for different relative intensities between the two drivers. Motivated by these experimental results, we focus on the theoretical investigation of the underlying physical mechanism at the single-atom level, providing a detailed analytical and numerical analysis. We demonstrate the interplay of three fundamental mechanisms responsible for the different contrast between neighboring harmonic lines. Their identification offers clear insight into the underlying electronic dynamics and the possibilities to control the ellipticity of the generated attosecond pulses.

The first mechanism at play is based on the Fano-Bethe type propensity rules in one-photon recombination [40]. The second mechanism is traced to the Barth-Smirnova type propensity rules in strong-field ionization [41–43]. The third mechanism is based on the impact of the two driving laser fields on the continuum-continuum transitions, i.e., the electron dynamics between ionization and recombination. The interplay of these three effects links the observed spectral features to the specific aspects of the subcycle electronic dynamics.

In particular, we show that the suppression of the (3N + 2) harmonic lines can be controlled by varying the relative intensity between the two fields, thus providing relatively easy means for controlling the ellipticity of the attosecond bursts, as suggested in [39]. We confirm our theoretical analysis with the numerical solution of the time-dependent Schrödinger equation (TDSE) and measurements in neon.

Furthermore, as inferred in [39], we show and explain that the suppression is not restricted to systems with outer electrons in p orbitals. Our theory also predicts, and the full solution of the TDSE confirms, the possibility of significant suppression

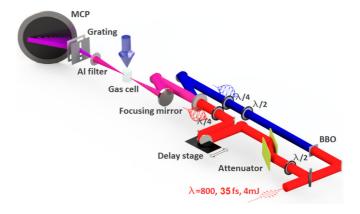


FIG. 2. Schematic of the optical setup. The BBO (beta-bariumborate) crystal was used for the generation of the second harmonic of the fundamental beam, MCP denoted the detector for the XUV radiation, and BS is the beam splitter.

of the (3N + 2) lines in the HHG spectrum of helium, starting with the ground s orbital. The suppression is achieved when the intensity of the fundamental field is sufficiently stronger than that of the second harmonic, eliminating the misconception that the desired suppression cannot happen when the electron is emitted from a spherically symmetric orbital, or that it requires the Cooper-minimum-like suppression of the recombination matrix elements [36].

II. EXPERIMENTAL OBSERVATION

We begin with the description of our experimental results, which motivate our numerical and analytical analysis. To generate the two-color bicircular fields, we have used a Ti:sapphire-based laser system with a single stage regenerative amplifier producing 35 fs pulses with up to 4 mJ energy and central wavelength of \sim 795 nm at 1 kHz repetition rate. The carrier-envelope phase (CEP) of the pulses was not locked. The laser beam was directed into the optical setup shown in Fig. 2.

The original beam was split into two beams with 60:40 ratio. The first, stronger, beam was sent onto a 0.1 mm thin BBO crystal to generate the second harmonic at \sim 400 nm with the pulse energy of up to 0.8 mJ. The second, weaker, beam remained at a fundamental wavelength of 795 nm with the pulse energy up to 1.5 mJ. The intensity was estimated based on comparing the observed cutoff position with the theoretical calculations (which is a standard approach; see, e.g. [18]), and taking into account the ratio of pulse energies and focusing conditions for the two incident driving fields. We could also tune smoothly the energies of the two pulses and their ratio by changing the original input pulse energy and by using a reflective attenuator (Fig. 2). Both beams were passed through achromatic broadband $\lambda/2$ and $\lambda/4$ waveplates, yielding nearly circular polarization ($\varepsilon \simeq 0.97$) for both the fundamental ("red") and its counter-rotating secondharmonic ("blue") beams. The beams were carefully combined in collinear geometry and focused with a single Ag mirror at f/100 into a 2-mm-long gas cell containing target gas. The cell was initially sealed with a metal foil, which was burned through by the laser beam at the start of the experiment. The resulting opening $d_0 \cong 60 \ \mu \text{m}$ was similar to the focal spot

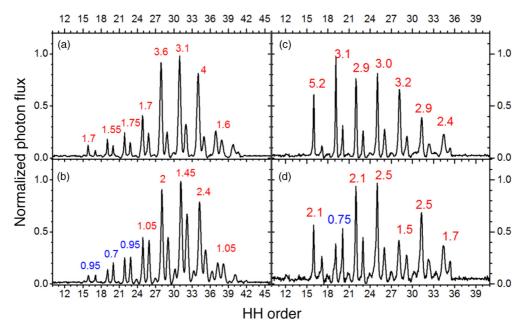


FIG. 3. High harmonics generated in neon for different peak intensities of the fundamental I_{ω} and second harmonic $I_{2\omega}$ circularly polarized fields. Intensities (in units of 10^{14} W/cm²) are estimated from the pulse energy, focal diameter, and the cutoff position: (a) $I_{\omega} \sim 4$, $I_{2\omega} \sim 3.3$; (b) $I_{\omega} \sim 2.6$, $I_{2\omega} \sim 3.3$; (c) $I_{\omega} \sim 2.5$, $I_{2\omega} \sim 2.0$; (d) $I_{\omega} \sim 1.6$, $I_{2\omega} \sim 2.0$. The ratios of the intensities of the adjacent harmonics with orders 3N+1 and 3N+2 are marked in the figures near to the corresponding harmonic pairs.

size, allowing us to keep the gas pressure inside the cell at ≈ 20 –40 mbar and the vacuum inside the interaction chamber at the level of $P_{\text{rest}} \approx 10^{-3}$ mbar. After passing the gas cell, the driving beams were blocked by a 200-nm-thick aluminum foil. The transmitted XUV radiation was directed towards the XUV spectrometer (for details, see [30]).

The XUV spectra generated in neon in Fig. 3 show the characteristic structure for the bicircular scheme. The harmonics with order 3N are strongly suppressed. The allowed 3N + 1 and 3N + 2 harmonics are nearly circularly polarized (see results of Refs. [22,27]), rotating in the same direction as the red and the blue beams, respectively.

Figures 3 and 4 show spectra of the XUV radiation generated for different relative intensities of the fundamental and the second harmonic fields. The driving fields were changed in such a way that the intensity of one color was fixed and the intensity of the other was varied. In Fig. 3 we have varied the intensity of the fundamental field, while in Fig. 4 the intensity of the second harmonic field was varied. In addition, the left and right columns in Figs. 3 and 4 present results for different values of the fixed fields.

Figures 3 and 4 demonstrate the dependence of the relative intensities of the 3N+1 and 3N+2 harmonics on the ratio between I_{ω} and $I_{2\omega}$. The increase of one driving field intensity relative to the other, independent of the total field strength, appears to enhance the generation of high harmonics corotating with it. Therefore, increasing the intensity of, e.g., the red field will change the ratio between the neighboring harmonics and will affect the overall chirality of the generated pulse train. Note that, if the ratio $I_{\omega}/I_{2\omega}$ is kept constant and both intensities are changed, the overall harmonic asymmetry is not affected, but the harmonic cutoff scales with the total intensity.

We also noted that when the intensity of the blue beam becomes larger than the intensity of the red beam, the forbidden 3N harmonics become more prominent in the spectrum [see Figs. 3(b), 3(d) and 4(b), 4(d)]. We discuss this phenomenon in detail in a separate publication, showing that it is related to the breaking of the dynamical symmetry in the system due to excitation of Rydberg states by the strong blue driver [30].

These experimental results demonstrate a simple practical way of controlling the degree of circularity of the attosecond pulse train by favoring harmonics with particular helicity, via changing the intensity ratio of the two driving fields. Of course, macroscopic effects may play a role in these findings. For example, the macroscopic effects can include the contribution of the free electrons to phase matching, and the number of such electrons changes with increasing intensity of the two driving fields. Nevertheless, the experimental trends in Figs. 3 and 4 remain consistent across a broad range of intensities of both fields. Below we show that the dominant role of the 3N + 1harmonics co-rotating with the red driver and the control over the ratio of 3N + 1 vs 3N + 2 harmonics originate already at the single-atom (microscopic) level, in accord with the results of [26,30]. We provide detailed analysis of this control and its physical origins using a theoretical model based on the strong field approximation (SFA), and then use time-dependent Schrödinger equation calculations to further corroborate our predictions.

III. METHOD

We consider a single active electron moving in the binding potential V_0 and interacting with the light field in the dipole approximation. In length gauge,

$$H = \frac{\hat{\mathbf{p}}^2}{2} + \hat{\mathbf{r}} \cdot \mathbf{E}(t) + V_0(\hat{\mathbf{r}}), \tag{1}$$

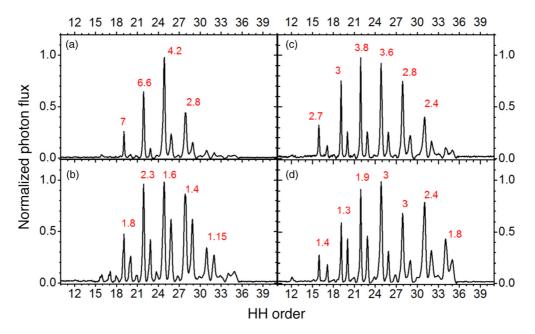


FIG. 4. High harmonics generated in neon for different peak intensities of the fundamental I_{ω} and second harmonic $I_{2\omega}$ fields. Intensities, in units of 10^{14} W/cm², were estimated from the pulse energy, focal diameter, and the cutoff position: (a) $I_{\omega} \sim 1.2$, $I_{2\omega} \sim 0.7$; (b) $I_{\omega} \sim 1.2$, $I_{2\omega} \sim 2.8$; (c) $I_{\omega} \sim 1.7$, $I_{2\omega} \sim 0.9$; (d) $I_{\omega} \sim 1.7$, $I_{2\omega} \sim 2.0$. The ratios of the intensities of the adjacent harmonics with orders 3N + 1 and 3N + 2 are marked in the figure near the corresponding harmonic pairs.

where $\hat{\mathbf{p}}$ is the momentum operator, $\hat{\mathbf{r}}$ is the position operator, and $\mathbf{E}(t)$ is the time-dependent electromagnetic field for the bicircular configuration, defined by

$$\mathbf{E}(t) = \operatorname{Re}\{-F_{\omega}e^{-i\omega t}\hat{\mathbf{e}}_{+} + F_{2\omega}e^{-i2\omega t}\hat{\mathbf{e}}_{-}\}. \tag{2}$$

We use atomic units unless otherwise stated. In the above, F_{ω} and $F_{2\omega}$ are the electric-field strengths of the fundamental and second harmonic fields, respectively, ω is the frequency of the fundamental field, and $\hat{\mathbf{e}}_{\pm} = \mp (\hat{\mathbf{e}}_x \pm i \, \hat{\mathbf{e}}_y)/\sqrt{2}$ are the unit vectors describing the polarization state of the driving fields. The unit vector $\hat{\mathbf{e}}_+$ corresponds to light rotating in the counterclockwise direction, while the unit vector $\hat{\mathbf{e}}_-$ corresponds to light rotating in the clockwise direction. In the following, we will use the same notation to indicate the polarization of the generated high harmonics and will depict the $\hat{\mathbf{e}}_\pm$ component as red $(\hat{\mathbf{e}}_+)$ and blue $(\hat{\mathbf{e}}_-)$ lines in the figures.

For the analytical description of the problem, we begin by assuming the core potential $V_0(\hat{\mathbf{r}})$ in Eq. (1) to be short range, defined by the conditions

$$H_0|\psi_{\ell m}\rangle = -I_p|\psi_{\ell m}\rangle,$$

$$\langle \mathbf{r}|\psi_{\ell m}\rangle = C_{\kappa\ell} \,\kappa^{3/2} (\kappa r)^{-1} e^{-\kappa r} Y_{\ell m}(\theta_r, \phi_r).$$
(3)

This approach, based on explicitly using the asymptotic behavior of the ground-state wave function outside the range of the potential (rather than the potential itself), is common to many analytical treatments [41,42,44–46], particularly with the effective range method developed by Frolov, Manakov, and Starace [45,46] and used extensively to study high harmonic generation [47–49]. With this procedure we neglect atomspecific features such as Cooper minima or Fano resonances, which can play a role for some atoms in some energy regions [36]. In the above, $\psi_{\ell m}$ is the ground state of the atomic system, with ionization potential I_p , $\kappa = \sqrt{2I_p}$, $Y_{\ell m}$ is the spherical

harmonic of angular and magnetic quantum number ℓ and m, respectively, and $C_{\kappa\ell}$ is a dimensionless real constant; its exact value is not relevant for our purposes.

The induced temporal dipole, proportional to the response of an individual atom or molecule to the laser field, can be written as [50]

$$\langle \Psi(t) | \hat{\mathbf{r}} | \Psi(t) \rangle = \sum_{m=-\ell}^{\ell} -i \int_{t_0}^{t} dt' e^{i I_p(t'-t)} \int d\mathbf{p} \, e^{-i S_V(\mathbf{p}, t, t')}$$

$$\times \mathbf{d}_{\ell m}^{(\text{rec})} [\mathbf{p} + \mathbf{A}(t)] \Upsilon_{\ell m} [\mathbf{p} + \mathbf{A}(t')] + \text{c.c.}, (4)$$

where "c.c." denotes the complex conjugate of the preceding expression, the sum over m takes into account the m degeneracy of the ground state, \mathbf{p} is the canonical momentum, and $\mathbf{A}(t) = -\int dt \, \mathbf{E}(t)$ is the vector potential of the field given by Eq. (2). The Volkov phase $S_V(\mathbf{p},t,t')$ is

$$S_V(\mathbf{p},t,t') = \frac{1}{2} \int_{t'}^t d\tau \left[\mathbf{p} + \mathbf{A}(\tau) \right]^2$$
 (5)

and the (scalar) ionization factor is

$$\Upsilon_{\ell m}[\mathbf{p} + \mathbf{A}(t')] = \left[\frac{[\mathbf{p} + \mathbf{A}(t')]^2}{2} + I_p\right] \langle \mathbf{p} + \mathbf{A}(t') | \psi_{\ell m} \rangle.$$
(6)

For the plane-wave continuum, the plane waves are defined as

$$\langle \mathbf{r} | \mathbf{p} + \mathbf{A}(t) \rangle = \frac{e^{i\mathbf{v}(t)\cdot\mathbf{r}}}{(2\pi)^{3/2}},$$
 (7)

where $\mathbf{v}(t) = \mathbf{p} + \mathbf{A}(t)$ is the kinetic momentum. The recombination dipoles are computed for the instantaneous electron velocity at the moment of emission t,

$$\mathbf{d}_{\ell m}^{(\text{rec})}[\mathbf{p} + \mathbf{A}(t)] = \langle \psi_{\ell m} | \hat{\mathbf{r}} | \mathbf{p} + \mathbf{A}(t) \rangle. \tag{8}$$

We note that this theoretical approach is well adapted to incorporate exact recombination dipoles (see, e.g. [48,49]). This expression for the induced dipole is a direct extension of the Perelomov, Popov, and Terent'ev approach to high harmonic generation [50]. The expression includes only the processes where the electron recombines to the same orbital from which it was ionized and assumes that there is no permanent dipole in the ground state.

The HHG spectrum at the frequency $N\omega$ is proportional to the induced frequency dipole, $\mathbf{D}(N\omega)$,

$$I(N\omega) \propto (N\omega)^4 |\mathbf{D}(N\omega)|^2,$$

$$\mathbf{D}(N\omega) = \int dt \langle \Psi(t) | \hat{\mathbf{r}} | \Psi(t) \rangle e^{iN\omega t}.$$
(9)

The five integrals in Eq. (9) can be computed using the saddle-point method (see, e.g. [50]). Amongst all possible electron trajectories, the saddle-point method selects those that contribute the most to the fivefold integral in Eq. (9), termed "quantum trajectories" (see, e.g. [51]) and reflecting purely quantum under-the-barrier motion [52] of the otherwise classical trajectories, with complex times of ionization and recombination. This leads to complex ionization and recombination velocities and to complex ionization and recombination angles, which would be crucial to the analysis of the high harmonic spectrum and, especially, to the contrast of the adjacent harmonic lines with opposite helicity.

Applying the saddle-point method, we can express the induced frequency dipole as

$$\mathbf{D}(N\omega) = \sum_{j}^{n_s} \mathbf{D}^{(j)}(N\omega),$$

$$\mathbf{D}^{(j)}(N\omega) \approx \sum_{m=-\ell}^{\ell} \mathbf{d}_{\ell m}^{(\text{rec})} \left[\mathbf{p}_s^{(j)} + \mathbf{A}(t_r^{(j)}) \right]$$

$$\times e^{-iS(\mathbf{p}_s^{(j)}, t_r^{(j)}, t_i^{(j)})} \Upsilon_{\ell m} \left[\mathbf{p}_s^{(j)} + \mathbf{A}(t_i^{(j)}) \right] e^{iN\omega t_r^{(j)}}. (10)$$

The quantities \mathbf{p}_s , t_r , and t_i are the complex-valued saddle-point solutions for the momentum, and times of recombination and ionization, respectively. The sum of index j runs over all the n_s stationary points. We will consider only those stationary points that correspond to the so-called short trajectories, which we can easily identify [21]. In this case, $n_s = 3M$, where M is the number of laser cycles.

Let us consider the contribution of a single laser cycle. The spectral intensity is the coherent sum of the three consecutive bursts. Since we are interested in the contrast between the two helicities of the emitted light $(\hat{\mathbf{e}}_+ \text{ and } \hat{\mathbf{e}}_-)$ in the spectra, it is useful to separate the intensity of the emitted light into the two helical components,

$$I(N\omega) = I_{+}(N\omega) + I_{-}(N\omega), \tag{11}$$

where

$$I_{\pm}(N\omega) = |A_{\pm}^{(1)}e^{i\phi_{\pm}^{(1)}} + A_{\pm}^{(2)}e^{i\phi_{\pm}^{(2)}} + A_{\pm}^{(3)}e^{i\phi_{\pm}^{(3)}}|^{2},$$
 (12)

with the amplitude and phase of the bursts corresponding to the amplitude and the phase of the frequency dipole in Eq. (10),

$$A_{\pm}^{(j)} = |D_{\pm}^{(j)}(N\omega)|, \quad \phi_{\pm}^{(j)} = \arg[D_{\pm}^{(j)}(N\omega)].$$
 (13)

Above, D_{\pm} is the $\hat{\mathbf{e}}_{\pm}$ helical component of the dipole vector \mathbf{D} , i.e., $D_{\pm} = \mathbf{D} \cdot \hat{\mathbf{e}}_{\pm}^*$, with $\hat{\mathbf{e}}_{\pm} \cdot \hat{\mathbf{e}}_{\pm}^* = 1$. In the long pulse limit, we have $A_{\pm}^{(1)} = A_{\pm}^{(2)} = A_{\pm}^{(3)} \equiv A_{\pm}$ and $\phi_{\pm}^{(j)} - \phi_{\pm}^{(j-1)} = (2\pi/3)(N \mp 1)$ (see Appendix \mathbf{D}), so that the intensity can be written as

$$I_{\pm}(N\omega) = A_{\pm}^{2} \left| 1 + 2 \cos \left[\frac{2\pi}{3} (N \mp 1) \right] \right|^{2}.$$
 (14)

The second term on the right-hand side of (14) is responsible for the well-known selection rules, i.e., harmonics $3N \pm 1$ are favored for $\hat{\mathbf{e}}_{\pm}$ and 3N harmonics are suppressed. The relative strengths of the harmonic lines, however, and thus the contrast between the (3N+1) and (3N-1) lines is contained exclusively in the amplitude A_+ .

According to Eq. (10), we may write

$$A_{\pm}^{2} = e^{2(\operatorname{Im}\{S\} - N\omega \operatorname{Im}\{t_{r}\})}$$

$$\times \left| \sum_{m=-\ell}^{\ell} \left(d_{\ell m, \pm}^{(\operatorname{rec})}[\mathbf{p}_{s} + \mathbf{A}(t_{r})] \right) \Upsilon_{\ell m}[\mathbf{p}_{s} + \mathbf{A}(t_{i})] \right|^{2}, \quad (15)$$

where $\operatorname{Im}\{x\}$ denotes the imaginary part of x and we have assumed that both the imaginary return time and the action do not change considerably from burst to burst, i.e., $\operatorname{Im}\{t_r^{(1)}\} \approx \operatorname{Im}\{t_r^{(2)}\} \approx \operatorname{Im}\{t_r^{(3)}\} \approx \operatorname{Im}\{t_r\}$ and $\operatorname{Im}\{S^{(1)}\} \approx \operatorname{Im}\{S^{(2)}\} \approx \operatorname{Im}\{S^{(3)}\} \approx \operatorname{Im}\{S\}$. Since $\exp[2(\operatorname{Im}\{S\} - N\omega \operatorname{Im}\{t_r\}]]$ is a scalar independent of m, it does not influence the contrast between the two helical components in the spectrum and we can ignore it for our purposes.

IV. PROPENSITY RULES IN TWO-COLOR HHG

In this section we will analyze how and why the strength of harmonic lines varies across the HHG spectrum of atoms emitting *s*-shell and *p*-shell electrons in the bicircular scheme. In particular, we will concentrate on two systems: helium and neon, whose HHG spectrum obtained by solving numerically the TDSE is shown in Fig. 1.

To solve the TDSE we used the code described in [53]. To simulate the neon atom, we used the 3D single-active electron pseudopotential given in [54], with a value of $z_{\rm eff}=10$ at the point r=0. To eliminate the contribution of long trajectories and bound states, and save computational time, we have used a small radial box of 70 a.u. The total number of points was nr=2500; we used a uniform grid spacing of 0.02 a.u. and a complex boundary absorber at 50 a.u. The time grid had a spacing of dt=0.0025 a.u. and the maximum angular momenta included in the expansion was $\ell_{\rm max}=60$. All the discretization parameters have been checked for convergence.

Equation (15) has now reduced our problem to the study of the ionization and recombination matrix elements within a single burst. Let us focus first on the one-photon recombination matrix element. To begin with, we point out one key aspect, which distinguishes one photon recombination in high harmonic emission from standard one-photon ionization and/or recombination process. This key aspect is often overlooked and implicitly ignored when one states that the recombination step in HHG is the (time-reversed) analog of one-photon ionization. The difference, however, is clear: in HHG recombination is conditioned on the electron return to the parent core and

therefore carries in it the imprint of the ionization step. In the quantum trajectory analysis, this imprint is encoded into the complex-valued recombination time t_r and (in general) the complex-valued velocity $\mathbf{v}(t_r)$. Below we will denote the real and imaginary parts of these and associated quantities with one or two "primes" correspondingly, e.g., $\text{Re}\{t_r\} = t_r'$, $\text{Im}\{t_r\} = t_r''$, etc.

For the one-photon dipole transition we apply the Wigner-Eckhart theorem and separate the radial part from the angular part:

$$d_{\ell m,\pm}^{(\text{rec})}[\mathbf{p}_{s} + \mathbf{A}(t_{r})] \equiv \langle \psi_{\ell m} | \hat{\mathbf{d}} \cdot \hat{\mathbf{e}}_{\pm}^{*} | \mathbf{v}(t_{r}) \rangle$$

$$= \sum_{\ell'=0}^{\infty} \mathcal{R}_{v(t_{r}),\ell'} e^{-i(m\pm 1)\phi'_{v(t_{r})}} \begin{pmatrix} \ell & \ell' & 1\\ -m & (m\pm 1) & \mp 1 \end{pmatrix}$$

$$\times \sqrt{\frac{(\ell' - m \mp 1)!}{(\ell' + m \pm 1)!}} P_{\ell'}^{(m\pm 1)} \left(\cos(\theta_{v(t_{r})}) \right) e^{(m\pm 1)\phi''_{v(t_{r})}}, \quad (16)$$

where $\mathbf{v}(t_r) = \mathbf{p}_s + \mathbf{A}(t_r)$ is the saddle-point complex velocity vector at the time of recombination, whose radial, polar, and azimuthal components are $v(t_r)$, $\theta_{v(t_r)}$, and $\phi_{v(t_r)}$, respectively. As pointed out above, the prime and double prime superscripts indicate, respectively, the real and imaginary parts of the corresponding quantity.

The radial factor $\mathcal{R}_{v(t_r),\ell'}$ is the same for $\hat{\mathbf{e}}_{\pm}$ and does not depend on the initial quantum number m, and hence does not influence the contrast between the lines (see Appendix C). The sum of index ℓ' runs over all angular momenta of the continuum states from which one-photon recombination occurs. Finally, the 3j coefficients reflect standard angular momentum algebra for one-photon transitions. For recombination to an s state, the 3j coefficients are zero unless $\ell'=1$, while for recombination to a p state, they are zero unless $\ell'=0$ or $\ell'=2$. The z component of the recombination velocity is negligible for collinear bicircular fields, so that $\cos\theta_{v(t_r)}=0$.

The difference between the recombination matrix element in Eq. (16) and its inverse photoionization counter-part is now clear. The inverse photoabsorption process occurs from a continuum state with real velocity. The recombination process, due to the fact that the electron is conditioned to return to the same place and state from which it was ionized, occurs from a continuum state with complex velocity vector—only its square has to be real valued.

This difference leads to the appearance of the last exponential factor in Eq. (16), which is absent in the inverse photoabsorption process. We may thus identify two separate contributions to the recombination dipole. The first is common with the photoionization process, and includes all of the terms in Eq. (16) except for the last exponential. The second is due to the recombination condition, which is given by the last exponential. Each of these two contributions is governed by a different propensity rule.

A. Propensity 1: Fano-Bethe rule

The first contribution, i.e., the emission analog of the photoabsorption process, follows the propensity rule given by Fano and Bethe [40,55]. This states that in one-photon transitions, the helicity of the emitted photon preferentially corotates with

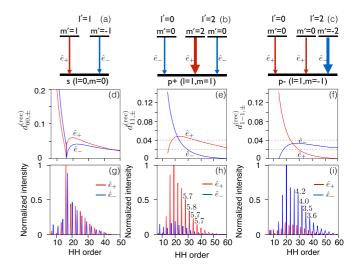


FIG. 5. Top: permitted dipole transitions in recombination to (a) s, (b) p+, and (c) p- states. Thick arrows indicate those favored by the Fano-Bethe propensity rule. Center: recombination matrix element [Eq. (16)], in arbitrary units, for final (d) s, (e) p+, and (f) p- states. Bottom: numerical TDSE high harmonic spectra when an electron is emitted from (g) s, (h) p+, and (i) p- states. Colored lines indicate the helicity of the emitted light: red for $\hat{\bf e}_+$ (counterclockwise) and blue for $\hat{\bf e}_+$ (clockwise).

the final state [see Figs. 5(b) and 5(c)]. Mathematically, in part this is a consequence of the fact that the 3j coefficient in Eq. (16) is larger when m = 1 for $\hat{\mathbf{e}}_+$ and when m = -1 for $\hat{\mathbf{e}}_-$. In Figs. 5(d)–5(f) we show the full recombination matrix element for the s ($\ell = 0, m = 0$), $p + (\ell = 1, m = 1)$, and $p - (\ell = 1, m = -1)$ states, as a function of the harmonic order. To extract the recombination velocities, we solved the corresponding saddle-point equations and used the results for the wave function in Eq. (3).

For helium [Figs. 5(d) and 5(g)], we considered an s-symmetric ground state with $I_p=0.9$ a.u. and driving fields of 12 fs duration and field strengths $F_{\omega}=F_{2\omega}=0.056$ a.u. For neon [Figs. 5(e), 5(f), 5(h), and 5(i)], we considered a p-symmetric ground state with $I_p=0.797$ a.u. and driving pulses of 12 fs duration and field strengths of $F_{\omega}=F_{2\omega}=0.074$ a.u. The pulse parameters and I_p were chosen to match those used in the TDSE calculations. As stated earlier, the $\hat{\bf e}_{\pm}$ component of the high harmonic lines are depicted as red or blue lines to highlight that they have the same helicity as that of the fundamental or second-harmonic driver.

For p+ [Figs. 5(b), 5(e), and 5(h)], the red lines corotate with the final state, while the blue lines counter-rotate with it. As predicted by the Fano-Bethe propensity rule, the red lines dominate the spectrum, with the exception of lower energies. The reason for the latter is simple. Recombination to a final state p+ happens from the two partial waves s and d, as shown in Fig. 5(b). While recombination from a d partial wave can occur to states corotating and counter-rotating with the field, recombination from an s wave is only possible to states counter-rotating with the field. At lower energies, the probability of the electron recombining from an s orbital is higher, and this is reflected in the spectrum as a dominance of the counter-rotating (blue) lines. For a final p- state [Figs. 5(c), 5(f), and 5(i)], the argument is analogous to that for p+, only

now the red lines counter-rotate with the final state, while the blue lines corotate with it.

It is worth noticing that the Fano-Bethe propensity rule cannot be modified by altering the parameters of the laser pulses, since it only depends on the modulus of the velocity, which is always fixed by the saddle-point solutions, i.e., $v(t_r) = \sqrt{2(N\omega - I_p)}$ [50].

B. Propensity 2: Recombination condition

If the recombination process in HHG were only the inverse of photoionization, then the recombination matrix elements for states p+ and p- would have been equivalent, up to the corresponding change in the helicity of the emitted light. A comparison of panels (e) and (f) in Fig. 5 shows that this is not the case. While the counter-rotating lines are the same (blue for p+ and red for p-), the corotating lines (red for p+ and blue for p-) are clearly different. The corotating line is more dominant in p+ than it is in p-.

This difference is the consequence of the recombination condition, i.e., the factor $e^{(m\pm 1)\phi_{v(t_r)}^n}$ in Eq. (16). This factor is equal to unity, and thus does not play a role for the counterrotating lines, but it enhances or suppresses the corotating lines, depending on $\phi_{v(t_r)}''$.

If $\phi''_{v(t_r)}$ is positive, which is the case for the vast majority of relevant harmonic orders and pulse parameters, then the red lines are enhanced in p+ while the blue lines are damped in p-. Intuitively, this propensity rule accounts for which photon is more likely to be absorbed or emitted in the free-free transition. As mentioned in [39], to lowest order, absorption of one extra photon of frequency ω (red) leads to the emission of the red harmonic line. Absorption of one extra photon of frequency 2ω (blue) leads to the emission of the blue harmonic line. For equal intensities of the ω and 2ω fields, the low-energy photons are more likely to cause free-free transitions [56], and thus the red lines dominate the harmonic spectrum. In contrast to the Fano-Bethe propensity rule described above, this propensity rule (in particular $\phi_{v(t_r)}''$) depends on the parameters of the driving fields through the saddle-point equations, and thus offers excellent opportunities for control.

When the electron is emitted from a single orbital, e.g., s, p+, or p-, the sum in Eq. (15) runs over one single index and the ionization term $\Upsilon_{\ell m}[\mathbf{p}_s + \mathbf{A}(t_i)]$ acts as a global factor. In this case, therefore, only the two mentioned above propensity rules will be responsible for the contrast between the harmonic lines in the spectrum. Indeed, in Figs. 5(g)-5(i) we show the solution of the TDSE for the case of ionization from s, p+and p- orbitals, respectively. The contrast between the two helicities (red and blue lines) follows nicely that predicted by our model, and thus follows the two simple propensity rules highlighted above. In particular, the numbers above the harmonic lines in Fig. 5(h) and Fig. 5(i) depict the ratio between the maxima of consecutive 3N + 1 and 3N + 2 lines in p+and between consecutive 3N + 2 and 3N + 1 lines in p-. The fact that red lines are stronger in p+ than blue lines are in p- is a manifestation of the recombination-conditioned propensity rule.

However, for all noble gas atoms except helium, the electron is emitted from both the p+ and p- orbitals. In this case, the ionization factor now plays a crucial role. In essence, it acts as

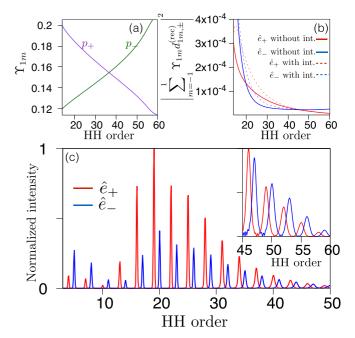


FIG. 6. (a) SFA-calculated ionization factor, in arbitrary units, of electrons emitted from the p+ (magenta) and p- (green) orbital of neon. (b) SFA-calculated square of the product of the ionization factor and recombination matrix element in the spectrum of neon separated in the two helical components (in arbitrary units), including (dashed lines) and neglecting (solid lines) the interference term in Eq. (17). (c) Spectrum of neon calculated solving the TDSE and separated in the two helical components [same as Fig. 1(b)].

an additional weighting factor for the contribution of p+ and p- electrons. Indeed, in this case the $\hat{\mathbf{e}}_{\pm}$ components of the intensity observed in the p spectrum are proportional to

$$I_{\pm} \propto |\langle \psi_{11} | d_{\pm} | \mathbf{v}(t_{r}) \rangle|^{2} |\Upsilon_{11} [\mathbf{v}(t_{i})]|^{2} + |\langle \psi_{1-1} | d_{\pm} | \mathbf{v}(t_{r}) \rangle|^{2} |\Upsilon_{1-1} [\mathbf{v}(t_{i})]|^{2} + F_{\text{int}} \cos[2(\phi'_{v(t_{i})} - \phi'_{v(t_{r})})] e^{\pm 2\phi''_{v(t_{r})}},$$
 (17)

where the first two terms above are the spectra observed when the electron is emitted from the p+ and p- orbitals, respectively, and the last term is the interference term which is different for each helicity component $\hat{\mathbf{e}}_{\pm}$. The phase of the interference term is composed of twice the phase difference between the ionization and recombination angles, and thus depends on the pulse parameters through the saddle-point equations. However, the influence of the interference term on the contrast between the lines, while slightly enhancing the red lines at low energies, is small compared to the preceding two contributions in Eq. (17) (see Fig. 6). It is therefore not a crucial term for our purposes and we will not comment on it further here. We note, nonetheless, that if one wishes to go beyond the plane-wave approximation and consider the real scattering states, then the interference term will contain the interplay between the scattering phases, which could lead to additional effects, as pointed out in [36].

Let us now concentrate on the first two terms in Eq. (17) and pose the following question: which orbital has the strongest ionization factor, p+ or p-? The answer is given by the third propensity rule, described below. Note that while the ionization

factor depends on the complex-valued ionization time t_i and the velocity $\mathbf{v}(t_i)$, it is also conditioned on the electron return to the parent core.

C. Propensity 3: Barth-Smirnova rule

Barth and Smirnova predicted that tunnel ionization driven by circular fields preferentially removes electrons from states counter-rotating with the field [41,42]. The same rule applies in the bicircular case [57,58], when the electron is required to return to the core. Incidentally, we note that this tunneling propensity rule is opposite to the photoionization one given by Fano-Bethe's propensity rule. In Fig. 6(a) we show the ionization factor, as a function of the harmonic number, for the electrons emitted from p+ and p- states, and for the same pulse parameters as those used in Fig. 5. The p+ state, which counter-rotates with respect to the total bicircular field, is dominant over the largest part of the relevant (more intense) part of the spectrum. The ionization factor, however, depends on the energy, and close to harmonic 40, emission from pstate becomes dominant. Again, we stress that this propensity rule depends on the pulse parameters, and can therefore be altered.

To calculate the contrast of red to blue lines in the p spectra within our model, we multiply the $\hat{\bf e}_+$ component of the recombination factor in the p+ spectrum [red line in Fig. 5(e)] by the p+ ionization factor [magenta line in Fig. 6(a)], and sum it with the $\hat{\mathbf{e}}_+$ component of the recombination factor in the pspectrum [red line in Fig. 5(f)] times the p- ionization factor [green line in Fig. 6(a)], according to Eq. (17). This gives us the amplitude of the $\hat{\mathbf{e}}_+$ (red) component in the p spectrum. Analogously, we obtain the $\hat{\bf e}_-$ (blue) component. The $\hat{\bf e}_+$ and $\hat{\bf e}_-$ amplitudes obtained this way are plotted in Fig. 6(b), as a function of the harmonic order. As commented earlier, Fig. 6(b) shows that neglecting the interference term [last term in Eq. (17)] gives essentially the same red and blue contrast, except at lower orders, where SFA is not accurate in the first place. The red and blue lines in the spectrum calculated from the TDSE [Fig. 6(c)] follows that predicted by the model [square of the solid red and blue lines in Fig. 6(b)] and coincides also with that predicted by the SFA in the rotating frame of reference [31].

We now have all the ingredients to understand the features in the HHG spectrum of neon, which we reproduce again in Fig. 6(c) for clarity. The spectrum shows a clear dominance of the red lines in the plateau region, as our model predicts. For increasing harmonic orders, the blue lines become comparable to the red lines and, near the cutoff region (\simeq HH45), the blue lines dominate. This feature is also well reproduced by our simple model. For low-energy harmonics the agreement is not so good, which is expected since the strong field approximation is not suited to treat below-threshold and close-to-threshold harmonics. Nonetheless, the higher contribution of blue lines at harmonics close to threshold, due to the stronger contribution of the s wave to the recombination process as highlighted earlier, is observed in both the spectrum and our model.

V. CONTROL OF THE HELICITY AND HARMONIC CONTRAST IN BICIRCULAR FIELDS

In the previous section we have outlined the three propensity rules responsible for the relevant features in the spectra. In this

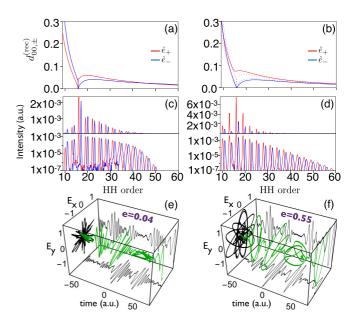


FIG. 7. Control of helicity of attosecond pulses generated in helium, using the bicircular driving field. Each of the two columns represents a different intensity ratio between the fundamental and the second harmonic: $I_{\omega}/I_{2\omega}=1$ for panels (a),(c) and $I_{\omega}/I_{2\omega}=2.25$ for panels (b),(d). Top row: SFA-calculated recombination matrix element for helium, as a function of harmonic order, in arbitrary units. Center row: TDSE-calculated spectrum of helium in linear (top) and logarithmic (bottom) scale. Bottom row: attosecond pulse trains (APT) generated by taking an energy window from harmonic 9 to harmonic 42. The degree of circularity e of the generated APTs is given (see text for details) and electric-field strengths are expressed in arbitrary units. The pulse parameters used in both the SFA and TDSE calculations are the same.

section we will use them to exert control over the ratio between the red and blue harmonic lines. As we have pointed out, the Fano-Bethe propensity rule cannot be altered by changing the pulse parameters, but we will now show how the recombination and ionization propensity rules modify the spectrum when the intensities of the two pulses are varied.

Let us first consider helium. As stated previously, in this case the ionization factor does not play any role in the contrast between the two helicities, so this contrast gives us direct access to the propensity rules in the recombination process. Figures 7(a) and 7(b) show the recombination factor of helium, $\langle \psi_{00} | \hat{\mathbf{d}} \cdot \hat{\mathbf{e}}_{+}^* | \mathbf{v}(t_r) \rangle$, for two intensity ratios between the fundamental and second harmonic, $I_{\omega}/I_{2\omega} = 1$ and 2.25, respectively. The duration of both driving pulses was set to 12 fs and the field strength of the second harmonic was always fixed at $F_{2\omega} = 0.056$ a.u. As the ratio increases, so does the contrast between the red and blue lines in both the recombination amplitude and in the HHG spectrum [Figs. 7(b) and 7(d)]. Higher red intensities translate into a higher number of red photons to be absorbed or emitted in the continuum and, consequently, to a stronger dominance of red harmonic lines in the spectrum.

In Figs. 7(e) and 7(f) we show the attosecond pulse train (APT) computed by inverse Fourier transform of the corresponding spectra, taking an energy window from harmonic 9

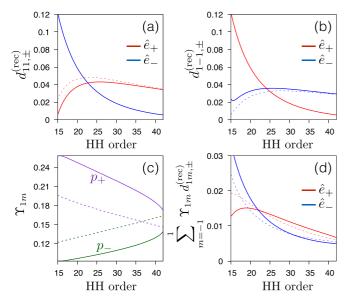


FIG. 8. Top row: SFA-calculated matrix elements for recombination to the (a) p+ and (b) p- orbitals of neon in arbitrary units. Bottom row: (c) SFA-calculated ionization factor of electrons emitted from the p+ (magenta) and p- (green) orbitals of neon and (d) SFA-calculated product of the ionization and recombination dipoles in the spectrum of neon, in arbitrary units. For all panels, solid red (blue) lines indicate emission of light with $\hat{\bf e}_+$ ($\hat{\bf e}_-$) helicity using a driver intensity ratio of $I_\omega/I_{2\omega}=0.5$. The dashed red (blue) lines indicate emission of light with $\hat{\bf e}_+$ ($\hat{\bf e}_-$) helicity using a driver intensity ratio of $I_\omega/I_{2\omega}=1$.

up to harmonic 42. When the intensity of the second harmonic driver is the same as that of the fundamental, there are three linearly polarized bursts per laser cycle ($T \approx 110$ a.u.). When the intensity of the fundamental is 2.25 that of the second harmonic [Fig. 7(f)], the bursts from the APT are highly elliptically polarized and rotating in the $\hat{\bf e}_+$ (counterclockwise) direction. We can calculate the degree of circularity of the generated attosecond pulse train by integrating the two helical components $E_{+} = \mp (E_{x} \pm i E_{y})/\sqrt{2}$ of the electric field over a temporal window, which we choose from $t_1 = -70$ a.u. to $t_2 = 70$ a.u. [same interval as that shown in Figs. 7(e) and 7(f)]. The degree of circularity can then be defined as $e = (|E_+|^2 - |E_-|^2)/(|E_+|^2 + |E_-|^2)$. Positive (negative) e will yield an attosecond pulse train elliptically polarized in the counterclockwise (clockwise) direction; e = 0 corresponds to a linearly polarized field, while |e| = 1 corresponds to a circularly polarized field. The value of e for the corresponding generated APT is shown in Figs. 7(e) and 7(f). We can control the polarization by simply changing the intensity of the red driving field with respect to the second harmonic. Note that, in this case, the generated bursts from the APT will always have the same helicity as the fundamental driving field, regardless of the frequency window applied to the spectrum to obtain them.

It is important to stress that the contrast between red and blue lines in helium appears because of the recombination-conditioned propensity rule, i.e., due to the fact that the trajectories propagate in complex space-time. If the electron trajectories moved in real-valued space-time, then no contrast would have been observed, and the APT generated in helium will always be linearly polarized, irrespective of the

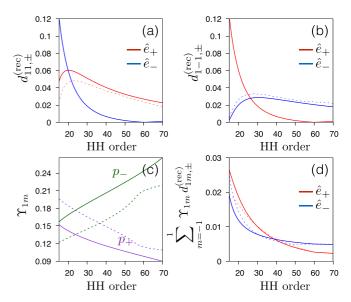


FIG. 9. Same as Fig. 8, but solid lines now indicate results using a driver intensity ratio of $I_{\omega}/I_{2\omega}=2$.

intensity ratio between the fundamental driver and the second harmonic.

We now turn our attention to neon. In Fig. 8 we show the recombination factor for p+ and p- [panels (a) and (b)], the ionization factor [panel (c)], and the resulting ratio between red and blue harmonic lines in the spectrum [panel (d)], when $I_{\omega}/I_{2\omega}=0.5$. In this case, there are more blue photons to be exchanged with the field during the free-free transitions, and hence the blue lines would be stronger than in the case of equal intensities. This is what we observe in the recombination factor, shown in Figs. 8(a) and 8(b) for the p+ and p- spectra, respectively. The ionization factor now favors the p+ spectrum throughout the below-threshold and plateau harmonics [Fig. 8(c)]. The p+ spectrum, however, has extended the dominance of the blue lines to higher harmonic orders. This, added to a decrease in the red:blue ratio, is detrimental to generating highly circular APT in the plateau. Additionally, increasing the intensity of the second harmonic increases the contribution of forbidden harmonics due to Rydberg state population, as seen in Figs. 3 and 4 and also in [30]. In the simulations, however, the small radial box used to reduce both computational time and the influence of long trajectories, eliminates the contribution of Rydberg orbits. The predicted contrast is given in Fig. 8(d), which is in good agreement with what the TDSE spectrum shows [Fig. 10(a)]. The pulses in this case were 12 fs long with a field strength of $F_{\omega} = F_{2\omega}/\sqrt{2} = 0.052$ a.u. The bursts of the APT generated in this case by filtering everything but the plateau region are close to linear [e = 0.23, Fig. 10(e)].

What happens when the fundamental field is stronger than the second harmonic? Again, we see a stronger dominance of red harmonic lines in the recombination factor in both the p+ and p- spectra due to the higher number of red photons exchanged in the continuum, as we saw in helium. As for the ionization factor, the p- spectrum now dominates over all the relevant harmonic region. The latter is responsible for an extended dominance of red lines to lower harmonic orders, as it

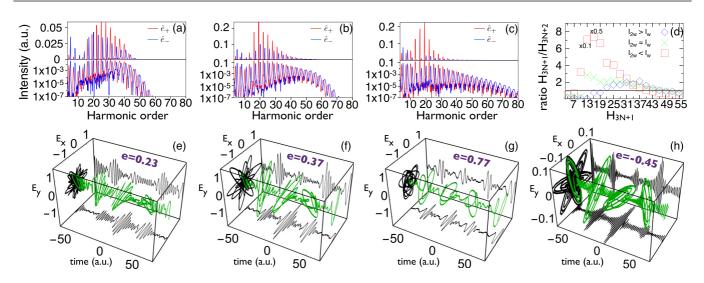


FIG. 10. Control of helicity of attosecond pulses generated in neon, using the bicircular driving field. TDSE-calculated HHG spectra in neon (left panels, with linear scale in the top and logarithmic scale in the bottom) and APT obtained from the spectra (right panels) for three different driver intensity ratios: (a), (e) $I_{\omega}/I_{2\omega} = 0.5$, (b),(f) $I_{\omega}/I_{2\omega} = 1$, and (c), (g), (h) $I_{\omega}/I_{2\omega} = 2$. The APTs in panels (e), (f), (g) were obtained by inverse Fourier transform of the plateau harmonics in panels (a), (b), (c), respectively, while the APT in panel (h) was obtained from the cutoff harmonics of panel (c). Electric-field strengths are expressed in arbitrary units. Panel (d) shows the ratio between the peak intensity of H_{3N+1} and H_{3N+2} for an $I_{\omega}/I_{2\omega}$ ratio of 0.5 (blue diamonds), 1 (green crosses), and 2 (red squares). A factor of 0.1 and 0.5 has been applied to H_{13}/H_{14} and H_{16}/H_{17} , respectively, in the case of $I_{\omega}/I_{2\omega} = 2$. The black line indicates equal intensity of H_{3N+1} and H_{3N+2} harmonics.

is illustrated in Fig. 9(d). The interplay between the ionization and recombination factors leads to a higher and more extended red-blue contrast in the spectrum in this case, as compared to the case of equal intensities. We observe this in both our model prediction [Fig. 9(d)] and the HHG spectrum [Fig. 10(c)], both computed with the same pulse parameters. In particular, we used a field strength of $F_{\omega} = \sqrt{2}F_{2\omega} = 0.1$ a.u. Therefore, just like in the case of helium, increasing the fundamental intensity with respect to the second harmonic dramatically increases the circularity of the APT generated with the plateau harmonics to the value e = 0.77, which we show in Fig. 10(g). In Fig. 10(d) we show the intensity ratio of the consecutive harmonics H_{3N+1}/H_{3N+2} for three different $I_{\omega}/I_{2\omega}$ ratios: 0.5 (blue diamonds), 1 (green crosses), and 2 (red squares). The ratio between consecutive harmonics clearly increases as the $I_{\omega}/I_{2\omega}$ ratio is increased for the most intense high harmonics. Furthermore, the higher total intensity in Fig. 10(c) allows us to look at the cutoff harmonics more clearly. The change from dominating red lines to blue lines in this region is evident, as also predicted by our model [see Fig. 9(d)]. In Fig. 10(h) we show the APT obtained from filtering out all harmonics below HH48. The bursts are remarkably elliptical and very clearly separated. The helicity of the bursts in this case is opposite to those obtained by filtering the plateau energies only, i.e., the bursts rotate in the $\hat{\mathbf{e}}_{-}$ direction, with a degree of circularity of e = -0.45. Therefore, we can obtain two elliptical attosecond pulse trains rotating in opposite directions within a single experiment just by spectral filtering.

VI. CONCLUSIONS

In conclusion, we have developed comprehensive analytical understanding of the electron dynamics occurring in the high harmonic generation process driven by two-color counterrotating fields and have linked these dynamics to the features observed in the experimental spectrum. We have identified the three propensity rules responsible for the contrast between the 3N+1 and 3N+2 harmonic lines in the HHG spectra of noble gas atoms, and demonstrated how these rules depend on the laser parameters and can be used to shape the polarization properties of the emitted attosecond pulses.

In particular, we were able to show that for atoms emitting from an *s*-shell (e.g., helium), the contrast between the two light helicities in the HHG spectra follows the one-photon recombination matrix element but still encodes the condition of the electron return. This offers one a unique opportunity to obtain the characteristics of what is, formally, recombination from a continuum state with complex-valued velocity.

For atoms emitting p electrons, information on the different ionization rates of p+ and p- electrons can similarly be obtained. This understanding of the dynamics demonstrates an easy to implement and efficient mechanism to control the polarization of the generated attosecond pulse trains. By increasing the ratio of intensities between the fundamental and second harmonic drivers, the contrast between the two adjacent harmonic lines in the spectrum dramatically increases, leading to more circular attosecond bursts. Moreover, we have showed that, for p states, when the fundamental field is stronger than the second harmonic, the APT generated from the plateau harmonics rotates with the fundamental driver, while that generated by the cutoff harmonics rotates with the second harmonic. In this way, we are able to obtain, from the same experiment, two circularly polarized APTs with opposite helicity. Contrary to the common view, we have shown that it is possible to generate highly chiral attosecond bursts by using initial s orbitals. Furthermore, we have linked this possibility to

the fact that the electron carries complex velocity at the time of recombination, and thus represents a measurable confirmation of the electron's propagation in complex space-time during the HHG process.

ACKNOWLEDGMENTS

A.J.G., N.Z., and M.I. acknowledge financial support from the DFG QUTIF Grant IV No. 152/6-1. D.A. and O.S. acknowledge support from the DFG Grant SM No. 292/2-3. E.P. acknowledges financial support from MINECO Grants FISICATEAMO (No. FIS2016-79508-P) and Severo Ochoa (No. SEV-2015-0522), Fundació Cellex, Generalitat de Catalunya (2014 SGR 874 and CERCA/Program), and ERC Grants EQuaM (FP7-ICT-2013-C No. 323714), QUIC (H2020-FETPROACT-2014 No. 641122), and OSYRIS (ERC-2013-ADG No. 339106).

APPENDIX A: RELATIONS BETWEEN VELOCITY COMPONENTS IN THE SADDLE-POINT METHOD

It is useful to derive some relations between the x and y components of the velocities that come out of the ionization and recombination saddle points, i.e.,

$$v(t_i)^2 = -2I_p, \quad v(t_r)^2 = 2(N\omega - I_p).$$
 (A1)

In general, both ionization and recombination velocities are complex.

$$v^{2} = v_{x}^{\prime 2} - v_{x}^{\prime \prime 2} + v_{y}^{\prime 2} - v_{y}^{\prime \prime 2} + 2i(v_{x}^{\prime}v_{x}^{\prime \prime} + v_{y}^{\prime}v_{y}^{\prime \prime}).$$
 (A2)

Using the saddle-point equations (A1), both ionization and recombination velocities have thus the fixed relation

$$v_x'v_x'' = -v_y'v_y''.$$
 (A3)

Using (A3), one can check that, since the square of the ionization velocity needs to be negative according to its saddle-point equation, necessarily

$$v_y''(t_i)^2 > v_x'(t_i)^2.$$
 (A4)

Analogously, for the recombination velocity we must have

$$v_y''(t_r)^2 > v_x'(t_r)^2$$
, for below threshold harmonics, $v_y''(t_r)^2 < v_x'(t_r)^2$, for above threshold harmonics, (A5) $v_y''(t_r)^2 = v_x'(t_r)^2$, at threshold.

Finally, we point out that, due to the symmetry of our field, both the real and imaginary parts of the ionization and recombination velocities of consecutive bursts transform following a clockwise rotation of $2\pi/3$, i.e.,

$$v_x^{(2)} = v_x^{(1)} \cos(2\pi/3) + v_y^{(1)} \sin(2\pi/3),$$

$$v_y^{(2)} = -v_x^{(1)} \sin(2\pi/3) + v_y^{(1)} \cos(2\pi/3).$$
(A6)

APPENDIX B: EXPRESSIONS FOR REAL AND IMAGINARY VELOCITY ANGLES

The saddle-point analysis yields complex kinetic momentum v_x and v_y , which can be expressed in complex spherical coordinates (v,θ,ϕ) . The module of this velocity is fixed by the saddle-point equation as highlighted in Appendix A. From (A1) one can see that the module of the velocity in ionization and in recombination for below threshold harmonics is imaginary. Hence, for ionization and recombination to below threshold harmonics, one can use the expansion of the trigonometric functions of complex arguments and write

$$\begin{aligned} v_z &= iv \cos \theta = 0, \\ v_x' &+ iv_x'' = iv \cos \phi = iv[\cos \phi' \cosh \phi'' - i \sin \phi' \sinh \phi''], \\ v_y' &+ iv_y'' = iv \sin \phi = iv[\sin \phi' \cosh \phi'' + i \cos \phi' \sinh \phi''], \end{aligned} \tag{B1}$$

where now v is real and is fixed up to a sign. We now identify the real and imaginary parts,

$$v_{\rm r}' = v \, \sin \theta' \, \sinh \phi'', \tag{B2}$$

$$v_r'' = v \cos \phi' \cosh \phi'', \tag{B3}$$

$$v_{v}' = -v \cos \phi' \sinh \phi'', \tag{B4}$$

$$v_{v}'' = v \sin \phi' \cosh \phi''. \tag{B5}$$

To obtain the real part of ϕ we can use, for example, Eqs. (B3) and (B5),

$$\phi' = \arctan \frac{v_y''}{v_x''},\tag{B6}$$

or (B2) and (B4),

$$\phi' = \arctan \frac{-v_x'}{v_y'}.$$
 (B7)

For the imaginary part we can use, for example, Eq. (B2) with (B5).

$$\phi'' = \operatorname{arctanh} \frac{v_x'}{v_y''},\tag{B8}$$

or (B4) with (B3),

$$\phi'' = \operatorname{arctanh} \frac{-v_y'}{v_x''}.$$
 (B9)

Indeed, the different choices are a consequence of the relations (A3). The apparent overdetermination is not such, since the choice we make here will automatically fix the sign of v, which is also an unknown. Indeed, if we use (B6) for the real part and (B8) for the imaginary part, then

$$v = -\frac{v_y'}{\cos \phi' \sinh \phi''},\tag{B10}$$

while if we use (B6) and (B9), then

$$v = \frac{v_x'}{\sin \phi' \sin \phi''}.$$
 (B11)

Equivalently, we can also choose (B7) and (B8), so that

$$v = \frac{v_x''}{\cos \phi' \cosh \phi''},\tag{B12}$$

or (B7) with (B9), and

$$v = \frac{v_y''}{\sin \phi' \cosh \phi''}.$$
 (B13)

Of course, regardless of the choice, the result is the same.

For recombination to above threshold harmonics, the same argument follows, only now the imaginary number need not be present in (B1) since the module of the velocity is already real. For example, one can choose to define

$$\phi' = \arctan \frac{v_y'}{v_x'},$$

$$\phi'' = \operatorname{arctanh} \frac{v_y''}{v_x'},$$

$$v = \frac{-v_x''}{\sin \phi' \sinh \phi''}.$$
(B14)

APPENDIX C: IONIZATION FACTOR AND RECOMBINATION MATRIX ELEMENT

The explicit steps of the derivation for the ionization factor can be found in [42], while that for the recombination matrix element directly follows from the same procedure. Here we just reproduce the results. For the ionization factor we have

$$\Upsilon[\mathbf{v}(t_i)] = \left[\frac{[\mathbf{v}(t_i)]^2}{2} + I_p\right] \langle \mathbf{v}(t_i) | \varphi_{\ell m}(\mathbf{r}) \rangle = \mathcal{F}_{\text{ion }} e^{im\phi_{v(t_i)}},$$
(C1)

where the factor

$$\mathcal{F}_{\text{ion}} = C_{\kappa\ell} \sqrt{\frac{2\kappa}{\pi}} \left(\frac{v(t_i)}{\kappa}\right)^{\ell} \sqrt{\frac{2\ell + 1(\ell - |m|)!}{4\pi(\ell + |m|)!}} P_{\ell}^{|m|}(\cos \theta_v)$$
(C2)

is independent on the sign of m. The complex angle $\phi_{v(t_i)}$ has the following real and imaginary components (see Appendix B):

$$\phi'_{v(t_i)} = \operatorname{atan} \frac{v''_{y}(t_i)}{v''_{y}(t_i)}, \quad \phi''_{v(t_i)} = \operatorname{atanh} \frac{v'_{x}(t_i)}{v''_{y}(t_i)}.$$
 (C3)

With the definition of the inverse hyperbolic tangent, $\operatorname{atanh}(z) = 1/2 \log \frac{1+z}{1-z}$, we can write the absolute value of the ionization step as

$$|\Upsilon[\mathbf{v}(t_i)]| = |\mathcal{F}_{\text{ion}}| \left(\frac{v_y''(t_i) + v_x'(t_i)}{v_y''(t_i) - v_x'(t_i)} \right)^{-m/2}, \quad (C4)$$

and the argument as

$$\arg \Upsilon[\mathbf{v}(t_i)] = m \operatorname{atan} \frac{v_y''(t_i)}{v_x''(t_i)} + \arg \mathcal{F}_{\text{ion}}.$$
 (C5)

Using (C4) along with (A3) and (A6), one finds that this absolute value is equal for the three bursts, so

$$\frac{\left|\Upsilon\left[\mathbf{v}\left(t_{i}^{(2)}\right)\right]\right|}{\left|\Upsilon\left[\mathbf{v}\left(t_{i}^{(1)}\right)\right]\right|} = \frac{\left|\Upsilon\left[\mathbf{v}\left(t_{i}^{(3)}\right)\right]\right|}{\left|\Upsilon\left[\mathbf{v}\left(t_{i}^{(1)}\right)\right]\right|} = 1.$$
 (C6)

We point out that \mathcal{F}_{ion} and, in particular, $v(t_i)$, does not change from burst to burst since it is fixed by the saddle-point equation $[v(t_i) = \sqrt{-2I_p}]$. Similarly, using (A6) along with the property atan $(\frac{x+y}{1-xy}) = atan(x) + atan(y)$, we obtain

$$\arg\left(\frac{\Upsilon\left[\mathbf{v}\left(t_{i}^{(2)}\right)\right]}{\Upsilon\left[\mathbf{v}\left(t_{i}^{(1)}\right)\right]}\right) = 2m\pi/3,$$

$$\arg\left(\frac{\Upsilon\left[\mathbf{v}\left(t_{i}^{(3)}\right)\right]}{\Upsilon\left[\mathbf{v}\left(t_{i}^{(1)}\right)\right]}\right) = 4m\pi/3.$$
(C7)

Following analogous steps as those given in [42], the recombination matrix element can be written as

$$d_{\ell m}^{\text{(rec)}}[\mathbf{v}(t_r)] = \mathcal{F}_{\text{rec}} e^{-i(m'\pm 1)\phi_{v(t_r)}},$$
(C8)

where, using Eq. (3),

$$\mathcal{F}_{\text{rec}} = (-1)^{m+1} \frac{C_{\kappa\ell}^*}{\sqrt{8\pi\kappa^5}} \sqrt{(2\ell+1)}$$

$$\times \sum_{\ell'=0}^{\infty} (2\ell'+1) \binom{\ell}{0} \binom{\ell'}{0} \binom{1}{0} \binom{\ell}{-m} \binom{\ell'}{(m\pm 1)} \binom{1}{+1}$$

$$\times \sqrt{\frac{(\ell'-m\mp 1)!}{(\ell'+m\pm 1)!}} P_{\ell'}^{(m\pm 1)} (\cos[\theta_{\nu}(t_r)]) \Gamma(\ell'+3)$$

$$\times \left(\frac{i\nu}{2\kappa}\right)^{\ell'} {}_{2}F_{1} \left(\ell'/2+3/2,\ell'/2+2;\ell'+3/2;-\frac{\nu^{2}}{\kappa^{2}}\right). \tag{C9}$$

The real and imaginary components of the complex angle $\phi_{v(t_r)}$ are

$$\phi'_{v(t_r)} = \operatorname{atan} \frac{v''_y(t_r)}{v''_x(t_r)}, \quad \phi''_{v(t_r)} = \operatorname{atanh} \frac{v'_x(t_r)}{v''_y(t_r)}, \quad (C10)$$

for below threshold harmonics, and

$$\phi'_{v(t_r)} = \operatorname{atan} \frac{-v''_x(t_r)}{v''_y(t_r)}, \quad \phi''_{v(t_r)} = \operatorname{atanh} \frac{v''_y(t_r)}{v'_y(t_r)}$$
(C11)

for above threshold harmonics. Again, with the definition of the inverse hyperbolic tangent we may write the absolute value of the recombination matrix element as

$$|d_{\ell,m,\pm}^{(\text{rec})}[\mathbf{v}(t_r)]| = |\mathcal{F}_{\text{rec}}|e^{(m'\pm 1)\phi_{v(t_r)}''}|$$

$$= |\mathcal{F}_{\text{rec}}|(-1)^{\gamma(m'\pm 1)/2}$$

$$\times \left(\frac{v_y''(t_r) + v_x'(t_r)}{v_y''(t_r) - v_x'(t_r)}\right)^{(m'\pm 1)/2}, \quad (C12)$$

where $\gamma=0$ or 1 depending on if we are considering below or above threshold harmonics. The expression for the argument of the recombination is

$$\arg\{d_{\ell,m,\pm}^{(\text{rec})}[\mathbf{v}(t_r)]\} = (m' \pm 1)\phi_{v(t_r)}' + \arg(\mathcal{F}_{\text{rec}}).$$
 (C13)

Using the same arguments as in the case of the ionization, the absolute value of the recombination for the three bursts is equivalent [the \mathcal{F}_{rec} factor does not change from burst to burst since it is independent on the velocity angle and $v(t_r)$ is fixed by the saddle-point equation],

$$\frac{\left|d_{\ell,m,\pm}^{(\text{rec})}[\mathbf{v}(t_r^{(2)})]\right|}{\left|d_{\ell,m,\pm}^{(\text{rec})}[\mathbf{v}(t_r^{(1)})]\right|} = \frac{\left|d_{\ell,m,\pm}^{(\text{rec})}[\mathbf{v}(t_r^{(3)})]\right|}{\left|d_{\ell,m,\pm}^{(\text{rec})}[\mathbf{v}(t_r^{(1)})]\right|} = 1.$$
(C14)

The relation between the argument of the recombination of consecutive bursts is

$$\arg \frac{d_{\ell,m,\pm}^{(\text{rec})} \left[\mathbf{v} \left(t_r^{(2)} \right) \right]}{d_{\ell,m,\pm}^{(\text{rec})} \left[\mathbf{v} \left(t_r^{(1)} \right) \right]} = -2(m \pm 1)\pi/3,$$

$$\arg \frac{d_{\ell,m,\pm}^{(\text{rec})} \left[\mathbf{v} \left(t_r^{(3)} \right) \right]}{d_{\ell,m,\pm}^{(\text{rec})} \left[\mathbf{v} \left(t_r^{(1)} \right) \right]} = -4(m \pm 1)\pi/3.$$
(C15)

APPENDIX D: DERIVATION OF RELATION BETWEEN AMPLITUDE AND PHASE OF DIFFERENT CONSECUTIVE BURSTS

In this section we wish to determine the relation between the amplitudes and phases in Eq. (13) of each of the three bursts in each cycle, in the monochromatic limit. We will assume that the initial and final state are given by a single orbital, i.e., the sum over m in Eq. (10) runs only over a single index. For a coherent superposition of two orbitals, like in the case of a p state (m = 1 and m = -1), the results are equivalent if one considers that the electron recombines to the same state from which it was ionized, i.e., $m_i = m_f$, where m_i and m_f are the initial and final magnetic quantum numbers, respectively. The amplitude and phase of the frequency dipole in Eq. (10) can be written as

$$\begin{aligned} |\mathbf{D}(N\omega)^{(j)}| &= e^{S''^{(j)} - N\omega t_r''^{(j)}} |\Upsilon[\mathbf{v}(t_i^{(j)})]| |\mathbf{d}_{\ell,m}^{(\text{rec})}[\mathbf{v}(t_r^{(j)})]|, \\ \arg\left[D_{\pm}(N\omega)^{(j)}\right] &= N\omega t_r'^{(j)} - S'^{(j)} + \arg\left\{\Upsilon[\mathbf{v}(t_i^{(j)})\right\} \\ &+ \arg\left\{d_{\ell,m,\pm}^{(\text{rec})}[\mathbf{v}(t_r^{(j)})]\right\}. \end{aligned} \tag{D1}$$

As always, the prime and double prime indicate the real and imaginary component of the corresponding quantity, respectively. In the monochromatic limit, the action is approximately the same for three consecutive bursts $S^{(1)} \approx S^{(2)} \approx S^{(3)}$. Due to the symmetry of the field, the real times of ionization and return are shifted by $2\pi/3$ between consecutive bursts, i.e., $t_r'^{(2)} \approx t_r'^{(1)} + \frac{2\pi}{3\omega}$ and $t_r'^{(3)} \approx t_r'^{(1)} + \frac{4\pi}{3\omega}$. The imaginary times of ionization and return, on the other hand, do not change for consecutive bursts, i.e., $t_r''^{(1)} \approx t_r''^{(2)} \approx t_r''^{(3)}$. Using the relations (C6), (C7), (C14), and (C15) derived in Appendix C, we get

$$|\mathbf{D}(N\omega)^{(1)}| = |\mathbf{D}(N\omega)^{(2)}| = |\mathbf{D}(N\omega)^{(3)}|,$$

$$\arg\left[D_{\pm}(N\omega)^{(j+1)}\right] - \arg\left[D_{\pm}(N\omega)^{(j)}\right] = \frac{2\pi}{3}(N \mp 1), \text{ (D2)}$$

as stated in the main text.

- [1] R. Cireasa, A. E. Boguslavskiy, B. Pons, M. C. H. Wong, D. Descamps, S. Petit, H. Ruf, N. Thire, A. Ferre, J. Suarez, J. Higuet, B. E. Schmidt, A. F. Alharbi, F. Legare, V. Blanchet, B. Fabre, S. Patchkovskii, O. Smirnova, Y. Mairesse, and V. R. Bhardwaj, Nat. Phys. 11, 654 (2015).
- [2] S. Beaulieu, A. Comby, B. Fabre, D. Descamps, A. Ferre, G. Garcia, R. Geneaux, F. Legare, L. Nahon, S. Petit, T. Ruchon, B. Pons, V. Blanchet, and Y. Mairesse, Faraday Discuss. **194**, 325 (2016).
- [3] C. Boeglin, E. Beaurepaire, V. Halte, V. Lopez-Flores, C. Stamm, N. Pontius, H. A. Durr, and J. Y. Bigot, Nature (London) 465, 458 (2010).
- [4] A. L. Cavalieri, N. Muller, T. Uphues, V. S. Yakovlev, A. Baltuska, B. Horvath, B. Schmidt, L. Blumel, R. Holzwarth, S. Hendel, M. Drescher, U. Kleineberg, P. M. Echenique, R. Kienberger, F. Krausz, and U. Heinzmann, Nature (London) 449, 1029 (2007).
- [5] C. E. Graves, A. H. Reid, T. Wang, B. Wu, S. de Jong, K. Vahaplar, I. Radu, D. P. Bernstein, M. Messerschmidt, L. Müller, R. Coffee, M. Bionta, S. W. Epp, R. Hartmann, N. Kimmel, G. Hauser, A. Hartmann, P. Holl, H. Gorke, J. H. Mentink, A. Tsukamoto, A. Fognini, J. J. Turner, W. F. Schlotter, D. Rolles, H. Soltau, L. Strüder, Y. Acremann, A. V. Kimel, A. Kirilyuk, T. Rasing, J. Stöhr, A. O. Scherz, and H. A. Dürr, Nat. Mater. 12, 293 (2013).
- [6] J.-Y. Bigot, Nat. Mater. 12, 283 (2013).
- [7] J.-Y. Bigot, M. Vomir, and E. Beaurepaire, Nat. Phys. 5, 515 (2009).

- [8] C. D. Stanciu, F. Hansteen, A. V. Kimel, A. Kirilyuk, A. Tsukamoto, A. Itoh, and T. Rasing, Phys. Rev. Lett. 99, 047601 (2007).
- [9] A. Kirilyuk, A. V. Kimel, and T. Rasing, Rev. Mod. Phys. 82, 2731 (2010).
- [10] I. Barth and O. Smirnova, J. Phys. B 47, 204020 (2014).
- [11] M. Hentschel, R. Kienberger, C. Spielmann, G. A. Reider, N. Milosevic, T. Brabec, P. Corkum, U. Heinzmann, M. Drescher, and F. Krausz, Nature (London) 414, 509 (2001).
- [12] M. Drescher, M. Hentschel, R. Kienberger, M. Uiberacker, V. Yakovlev, A. Scrinzi, T. Westerwalbesloh, U. Kleineberg, U. Heinzmann, and F. Krausz, Nature (London) 419, 803 (2002).
- [13] E. Goulielmakis, Z.-H. Loh, A. Wirth, R. Santra, N. Rohringer, V. S. Yakovlev, S. Zherebtsov, T. Pfeifer, A. M. Azzeer, M. F. Kling, S. R. Leone, and F. Krausz, Nature (London) 466, 739 (2010)
- [14] G. Sansone, F. Kelkensberg, J. F. Perez-Torres, F. Morales, M. F. Kling, W. Siu, O. Ghafur, P. Johnsson, M. Swoboda, E. Benedetti, F. Ferrari, F. Lepine, J. L. Sanz-Vicario, S. Zherebtsov, I. Znakovskaya, A. L'Huillier, M. Y. Ivanov, M. Nisoli, F. Martin, and M. J. J. Vrakking, Nature (London) 465, 763 (2010).
- [15] F. Krausz and M. Ivanov, Rev. Mod. Phys. 81, 163 (2009).
- [16] F. Calegari, D. Ayuso, A. Trabattoni, L. Belshaw, S. De Camillis, S. Anumula, F. Frassetto, L. Poletto, A. Palacios, P. Decleva, J. B. Greenwood, F. Martín, and M. Nisoli, Science 346, 336 (2014).

- [17] V. Gruson, L. Barreau, Á. Jiménez-Galan, F. Risoud, J. Caillat, A. Maquet, B. Carré, F. Lepetit, J. F. Hergott, T. Ruchon, L. Argenti, R. Taïeb, F. Martín, and P. Salières, Science 354, 734 (2016).
- [18] O. Smirnova, Y. Mairesse, S. Patchkovskii, N. Dudovich, D. Villeneuve, P. Corkum, and M. Y. Ivanov, Nature (London) 460, 972 (2009).
- [19] H. Eichmann, A. Egbert, S. Nolte, C. Momma, B. Wellegehausen, W. Becker, S. Long, and J. K. McIver, Phys. Rev. A 51, R3414 (1995).
- [20] T. Zuo and A. D. Bandrauk, J. Nonlinear Opt. Phys. Mater. 04, 533 (1995).
- [21] D. B. Milošević, W. Becker, and R. Kopold, Phys. Rev. A 61, 063403 (2000).
- [22] A. Fleischer, O. Kfir, T. Diskin, P. Sidorenko, and O. Cohen, Nat. Photon. 8, 543 (2014).
- [23] E. Pisanty, S. Sukiasyan, and M. Ivanov, Phys. Rev. A 90, 043829 (2014).
- [24] M. Ivanov and E. Pisanty, Nat. Photon. 8, 501 (2014).
- [25] D. D. Hickstein, F. J. Dollar, P. Grychtol, J. L. Ellis, R. Knut, C. Hernández-García, D. Zusin, C. Gentry, J. M. Shaw, T. Fan, K. M. Dorney, A. Becker, A. Jaroń-Becker, H. C. Kapteyn, M. M. Murnane, and C. G. Durfee, Nat. Photon. 9, 743 (2015).
- [26] L. Medišauskas, J. Wragg, H. van der Hart, and M. Y. Ivanov, Phys. Rev. Lett. 115, 153001 (2015).
- [27] O. Kfir, P. Grychtol, E. Turgut, R. Knut, D. Zusin, D. Popmintchev, T. Popmintchev, H. Nembach, J. M. Shaw, A. Fleischer, H. Kapteyn, M. Murnane, and O. Cohen, Nat. Photon. 9, 99 (2015).
- [28] A. Ferré, C. Handschin, M. Dumergue, F. Burgy, A. Comby, D. Descamps, B. Fabre, G. A. Garcia, R. Géneaux, L. Merceron, E. Mével, L. Nahon, S. Petit, B. Pons, D. Staedter, S. Weber, T. Ruchon, V. Blanchet, and Y. Mairesse, Nat. Photon. 9, 93 (2015).
- [29] T. Fan, P. Grychtol, R. Knut, C. Hernández-García, D. D. Hickstein, D. Zusin, C. Gentry, F. J. Dollar, C. A. Mancuso, C. W. Hogle, O. Kfir, D. Legut, K. Carva, J. L. Ellis, K. M. Dorney, C. Chen, O. G. Shpyrko, E. E. Fullerton, O. Cohen, P. M. Oppeneer, D. B. Milošević, A. Becker, A. A. Jaroń-Becker, T. Popmintchev, M. M. Murnane, and H. C. Kapteyn, Proc. Natl. Acad. Sci. USA 112, 14206 (2015).
- [30] Á. Jiménez-Galán, N. Zhavoronkov, M. Schloz, F. Morales, and M. Ivanov, Opt. Express 25, 22880 (2017).
- [31] E. Pisanty and Á. Jiménez-Galán, Phys. Rev. A 96, 063401 (2017).
- [32] O. Kfir, P. Grychtol, E. Turgut, R. Knut, D. Zusin, A. Fleischer, E. Bordo, T. Fan, D. Popmintchev, T. Popmintchev, H. Kapteyn, M. Murnane, and O. Cohen, J. Phys. B 49, 123501 (2016).
- [33] C. Hernández-García, C. G. Durfee, D. D. Hickstein, T. Popmintchev, A. Meier, M. M. Murnane, H. C. Kapteyn,

- I. J. Sola, A. Jaron-Becker, and A. Becker, Phys. Rev. A **93**, 043855 (2016).
- [34] A. Bandrauk, F. Mauger, and K.-J. Yuan, J. Phys. B **49**, 23LT01 (2016).
- [35] S. Odžak, E. Hasović, and D. B. Milošević, Phys. Rev. A 94, 033419 (2016).
- [36] D. Baykusheva, M. S. Ahsan, N. Lin, and H. J. Wörner, Phys. Rev. Lett. 116, 123001 (2016).
- [37] O. E. Alon, V. Averbukh, and N. Moiseyev, Phys. Rev. Lett. 80, 3743 (1998).
- [38] D. B. Milošević, Opt. Lett. 40, 2381 (2015).
- [39] K. M. Dorney, J. L. Ellis, C. Hernández-García, D. D. Hickstein, C. A. Mancuso, N. Brooks, T. Fan, G. Fan, D. Zusin, C. Gentry, P. Grychtol, H. C. Kapteyn, and M. M. Murnane, Phys. Rev. Lett. 119, 063201 (2017).
- [40] U. Fano, Phys. Rev. A 32, 617 (1985).
- [41] I. Barth and O. Smirnova, Phys. Rev. A 84, 063415 (2011).
- [42] I. Barth and O. Smirnova, Phys. Rev. A 87, 013433 (2013).
- [43] J. Kaushal, F. Morales, and O. Smirnova, Phys. Rev. A 92, 063405 (2015).
- [44] A. M. Perelomov, V. S. Popov, and M. V. Terent'ev, Zh. Eksp. Teor. Fiz. 50, 1393 (1966) [Sov. Phys. JETP 23, 924 (1966)].
- [45] M. V. Frolov, N. L. Manakov, E. A. Pronin, and A. F. Starace, Phys. Rev. Lett. 91, 053003 (2003).
- [46] M. V. Frolov, N. L. Manakov, and A. F. Starace, Phys. Rev. A 78, 063418 (2008).
- [47] M. V. Frolov, N. L. Manakov, T. S. Sarantseva, and A. F. Starace, J. Phys. B 42, 035601 (2009).
- [48] M. V. Frolov, N. L. Manakov, and A. F. Starace, Phys. Rev. Lett. 100, 173001 (2008).
- [49] M. V. Frolov, N. L. Manakov, A. A. Silaev, N. V. Vvedenskii, and A. F. Starace, Phys. Rev. A 83, 021405 (2011).
- [50] O. Smirnova and M. Ivanov, Multielectron high harmonic generation, in *Attosecond and XUV Physics* (Wiley-VCH Verlag GmbH & Co. KGaA, Berlin, 2014), Chap. 7, pp. 201–256.
- [51] P. Salières, B. Carré, L. Le Déroff, F. Grasbon, G. Paulus, H. Walther, R. Kopold, W. Becker, D. Milošević, A. Sanpera et al., Science 292, 902 (2001).
- [52] M. Lewenstein, P. Balcou, M. Y. Ivanov, A. L'Huillier, and P. B. Corkum, Phys. Rev. A 49, 2117 (1994).
- [53] S. Patchkovskii and H. G. Muller, Comput. Phys. Commun. 199, 153 (2016).
- [54] X. M. Tong and C. D. Lin, J. Phys. B 38, 2593 (2005).
- [55] H. Bethe and R. Jackiw, *Intermediate Quantum Mechanics*, 3rd ed. (Westview Press, USA, 1997), Chap. 11.
- [56] Á. Jiménez-Galán, F. Martín, and L. Argenti, Phys. Rev. A 93, 023429 (2016).
- [57] D. Ayuso, A. Jiménez-Galán, F. Morales, M. Ivanov, and O. Smirnova, New J. Phys. 19, 073007 (2017).
- [58] D. B. Milošević, Phys. Rev. A 93, 051402 (2016).